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Berndt Mueller (Duke University)

# PHENOMENOLOGY

### Probes of thermalization

- Particle p<sub>T</sub> spectra
- Hadron yields
  - Rafelski, Cleymans
- Collective flow
  - Ma, Hirano, Csernai, Borghini, Lacey, Grassi
- Heavy quarks (D, J/Ψ, etc.)
  - □ Nagle, Thews
- Jets, photons, and other hard probes
   Cole, Odyniec

### A few "obvious" remarks

- Matter created in a heavy ion collision cannot be *truly* thermal. Perfect thermalization requires either gravity or a static, confining box.
- If a QGP is formed, J/Ψ regeneration must occur, the question is how much. Any set of kinetic mechanisms will bring the number of J/Ψ closer to equilibrium.
- Deviations of data from predictions can be a sign of real physics (such as incomplete equilibration) or of the inadequacy of model assumptions.

### **3-D Hydrodynamics**

- 3-D relativistic hydro codes with sophisticated freeze-out modeling are becoming state of the art.
- Microscopic treatment of hadronic phase unavoidable.
- Need to understand range of admissible initial conditions, EOS, dependence on details of freeze-out treatment.
  - $\Box$  E.g.: How does v<sub>4</sub>/v<sub>2</sub><sup>2</sup> depend on  $\alpha_4/\alpha_2$  of initial shape?
- Hydro codes must be applied to systematic studies of AA phenomenology: flow, hadron spectra, imbedded hard probes (photons, jets, charm).
- How far can we go without QGP viscosity? Borghini's study suggests that Kn > 0 effects are still sizable at RHIC.
  - $\square$  Prediction and test: v<sub>2</sub> still grows at LHC !

# (MICROSCOPIC) TRANSPORT THEORY

### Parton transport

- Parton cascade with gg ↔ ggg and detailed balance equilibrates remarkably rapidly (Xu, Xu).
  - $\Box$  More efficient than  $\sigma_{tr}$  would naively suggest.
  - Do calculation in weak coupling limit to check whether it agrees with the "bottom-up" scenario. If not, why not ?

 $\mathbf{a}\mathbf{m}^2$ 

 Connection between shear viscosity and charm energy loss / flow (Teaney). Collisional Langevin equation for charm quark:

$$\frac{dp}{dt} = \eta_D p + \xi(t) \quad \text{with} \quad \eta_D = \frac{T}{MD}, \left\langle \xi_i(t)\xi_k(t') \right\rangle = \frac{2T^2}{3D} \delta_{ik}\delta(t-t')$$

$$pQCD: \quad D \approx \frac{6}{8\pi\alpha_s^2 T} \approx 6\frac{\eta}{sT} \quad \xrightarrow{\text{PHENIX}}_{\text{Data}} \quad D = \frac{3\cdots6}{2\pi T} \Leftrightarrow \frac{\eta}{s} = \frac{1\cdots2}{4\pi}$$

### Away-side jets



STAR result: di-jet suppression  $R_{AA} \approx I_{AA} \approx 0.23$ Corona effect  $R_{\rm AA} \sim \Delta R/R$ Halo effect Away-side jet  $I_{\scriptscriptstyle AA} = R_{\scriptscriptstyle AA}^{\rm dijet} / R_{\scriptscriptstyle AA} \sim \Delta R / R$ 

#### Jet correlations of secondaries



### Collective motion or heat ?

- What happens to the radiated energy?
- Extreme alternatives:
  - The radiation propagates (nearly) frictionless as quasiparticle modes, which may fragment further and eventually hadronize outside the dense medium.
  - The radiation gets absorbed in the medium and heats it locally. The radiated energy and momentum is shared with the absorbing medium.

• General case:  $f_{mach} + f_{heat} = 1$ .

### Extreme scenarios



### Mach cones in an evolving medium

- Fireball evolution model describing hadronic m<sub>t</sub> spectra, HBT, photon emission and R<sub>AA</sub> (T. Renk, Phys. Rev. C 70 (2004) 021903)
- Space-time position dependent jet energy deposition
- Fraction *f*<sub>mach</sub> of energy deposition into collective mode ("sound")
- Local speed of sound c<sub>s</sub> (and fireball thermodynamics) from lattice EOS
- Propagation of sound waves through evolving medium (incl. flow)
- Freeze-out using Cooper-Frye formula
- Monte-Carlo sampling using trigger conditions and acceptance cuts

### Comparison with data



forward jet momentum (colorless sound)

### Medium modified MLLA



#### Borghini & Wiedemann



# MECHANISMS OF THERMALIZATION

### Thermalization

- Matters of principle
  - □ Kovchegov
- Thermalization in (resummed) perturbation theory
  - Mueller, Borsanyi, Arrizabalaga
- Plasma instabilities
  - □ Moore, Manuel, Nara, Strickland, Romatschke
- Quarks
  - 🗆 Lappi
- Prethermalization
  - Berges, Kharzeev

### Modified BUS

New scaling solution does not change the parametric dependence of long the tail (depletion of hard partons):  $\tau_{th} \sim \alpha_s^{-13/5} Q_s^{-1}$  [ ~ 6  $Q_s^{-1} \sim 1$  fm/c for  $\alpha_s = 0.5$  !]

But what is the initial time?



Biro et al., PLB 283 (1992) 171, PRC 48 (1993) 1275

 $p_{T/L}$ 

### **Color instabilities**

Unstable modes occur generally due to the anisotropy of the momentum distribution:

$$\left\langle \left(\Delta p_x\right)^2 \right\rangle = \left\langle \left(\Delta p_y\right)^2 \right\rangle = Q_s^2 \gg \left\langle \left(\Delta p_z\right)^2 \right\rangle \quad \text{for} \quad \tau \gg \frac{1}{Q_s}$$



p<sub>v</sub>

Wavelength and growth rate of unstable modes can be calculated perturbatively:

 $k_z \sim gQ_s$ ,  $\gamma \sim gQ_s < k_z$ 

Moore, Strickland, Dumitru, Manuel

### Color instabilities (2)



## Color instabilities (3)

1. Do longitudinal instabilities persist, if "hard" particles are replaced by "hard" classical color fields? Maybe not (Romatschke, but see: W. Pöschl & BM, CPC 125 (2000) 282, PRD 60 (99) 114505).



FIG. 8. The transverse color electric field energy dansity w<sup>(j)</sup><sub>i</sub>(t, x<sub>j</sub>, x<sub>j</sub>) is shown for six selected snapshots lakes at the time steps t<sub>ime</sub>, t<sub>1000</sub>, t

# THE STRUCTURE OF THERMAL QCD MATTER

# The QGP near T<sub>c</sub>

- Perturbative analysis of QGP for T ≥ 2 3T<sub>c</sub> has been achieved by loop expansion around HTL quasiparticles (Blaizot, Iancu, Rebhan; Andersen, Braaten, Strickland). What happens nearer to T<sub>c</sub>?
- Are there quasiparticles with Γ < M ? If so, what are these quasiparticles? Is there short or long range order?
- Pair correlation function, e.g. g(r) = (y<sup>+</sup>(0)y(0)y<sup>+</sup>(r)y(r)) could show a small *peak* at r=0 for moderate coupling, because of attractive channels.

### The SUSY connection

#### • A challenge to physicists:

- □ What is the limit of perfection?
- Strong coupling limit of N = 4
   SYM (Starinets):

$$\frac{\eta}{s} = \frac{1}{4\pi} \left( 1 + \frac{135\zeta(3)}{8(2g^2N_c)^{3/2}} \right)$$

- □ Does limit apply to QCD ?
- No long-range structure formation in SYM (no large mass scale compared with *T*).



### QCD molecular dynamics?

$$\frac{dp_i^{\mu}}{d\tau} = gQ_i^a F^{a\mu\nu} \mathbf{v}_{i\nu} \quad \bigoplus \quad \frac{dQ_i^a}{d\tau} = gf_{abc} A^{a\mu} Q_i^c \mathbf{v}_{i\mu} \qquad \text{Wong}$$
$$\partial_{\mu} F^{a\mu\nu} = j^{a\nu} \equiv \int d\tau \sum_i gQ_i^a \mathbf{v}_i^{\nu} \delta\left(x - \xi_i(\tau)\right) \qquad \text{Heinz}$$

Reproduces HTL pert. th. at weak coupling (Kelly, Liu, Lucchesi, Manuel)

Lattice/particle formulation exists (Hu, BM, Moore - Dumitru, Nara)





"Well suited" for studies of equilibrium properties

(problems with transport)



### Participants of QCD-TH @ Vienna

