# NC Emergent Gravity & Fermions

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# Outline

To begin with ... an Introduction
Geometry & NC *U(N)* gauge theory
Scalars

Fermions

... then gravity emerges ...

Geometry & UV/IR mixing
Conclusion

Motivation
The Matrix Mov

Scalars

Fermions

### Emergent Gravity

Unimodular Me Seeley-de Witt

EH Action

JV/IR mixin

Conclusion

# Introduction

, GR & QM ?

Quantum fluctuations of space-time

Noncommutative space-time

$$[x_i, x_j] = i \,\theta_{ij}$$

Desire simple, intrinsic relation between NC & Gravity

Motivatio

he Matrix Model calars

### Emergent Gravity

Upon Quantization
Unimodular Metric
Seeley-de Witt

EH Action

UV/IR mixing

Conclusions

# NC gauge theory contains gravity

Motivation
The Matrix Mode

Fermion

Gravity
Upon Quantiza

Unimodular Met Seeley-de Witt coefficients

EH Actio

JV/IR mixin

Conclusion

# The Model

$$S_{YM}=- ext{Tr}ig[Y^a,Y^big]ig[Y^{a'},Y^{b'}ig]g_{aa'}g_{bb'}$$
  $Y^a\in L\left(\mathcal{H}
ight)$  ... matrices/operators;  $a=0,1,2,3;$   $ig[Y^a,Y^big]=i\, heta^{ab}$ 

- $\theta^{ab}$  not constant
- $Y^a$  interpreted as quantization of coordinate functions  $y^a$  on Poisson manifold  $(\mathcal{M}, \theta^{ab}(y))$  with Poisson structure  $\theta^{ab}(y)$
- This implies in the semi classical limit

$$[Y^a, \Psi] \sim i \, \theta^{ab}(y) \frac{\partial}{\partial y^b} \Psi$$

# Scalars

## **Effective Metric**

$$S [\Phi] = -\text{Tr} [Y^a, \Phi] [Y^b, \Phi] g_{ab}$$
$$\sim \int d^4 y \, \rho(y) G^{ab}(y) \partial_a \Phi(y) \partial_b \Phi(y)$$

where

$$G^{mn}(y) = \theta^{ma}(y)\theta^{nb}(y)g_{ab}$$

- Effective metric determined by Poisson sturcture  $\theta^{mn}$
- Φ couples to effective metric
- $\rho(y) = |G_{ab}|^{1/4}$
- $\theta^{mn} \rightarrow \text{vielbein}$

EH Actio

UV/IR mixin

Conclusion

# Fermions

# The Action

$$S = (2\pi)^2 \operatorname{Tr} \bar{\Psi} \gamma_a [Y^a, \Psi] \ \sim \int \mathrm{d}^4 y 
ho(y) \bar{\Psi} i \gamma_a heta^{ab}(y) \partial_b \Psi$$

The Dirac operator

$$\not\!\!D\Psi = \gamma_a [Y^a, \Psi] \sim i \gamma_a \theta^{ab}(y) \partial_b \Psi$$

★ no spin connection appears

Compare to standard covariant derivative

Motivation
The Matrix Model
Scalars

### Emergen Gravity

Upon Quantizatio Unimodular Metr Seeley-de Witt coefficients

EH Action

UV/IR mixii

Conclusion

# Spin connection?

# Yes, this is a strange feature

- Rotation of spin will be different in this model
- Holonomies will be different than in GR

### Nevertheless ...

... trajectory of fermion will follow geodesics ... Einstein-Hilbert action will be induced

# This is a reasonable action for fermions

... Opens possiblity for experimental signature?

Motivation
The Matrix Mode

Fermions

Gravity

Upon Quantizati

Unimodular N Seeley-de Wit coefficients

EH Actio

JV/IR mixin

Conclusion

# **Emergent Gravity**

How will the Einstein-Hilbert action join our game? No need of adding further terms to the action.

 $S_{EH}$  emerges automatically upon quantization. One-loop effective action  $\Gamma_{\Psi}$  will correspond to EH-action  $S_{EH}$ .

$$e^{-\Gamma_{\Psi}} = \int d\Psi d\bar{\Psi} e^{-S[\Psi,\bar{\Psi}]}$$

$$\Gamma_{\Psi} = -\frac{1}{2} \text{Tr} \log \not D^2$$

$$S_{
m square} = \left(2\pi^2
ight) {
m Tr} ar{\Psi} D\!\!\!/ \, {}^2 \Psi = \int d^4 y 
ho(y) ar{\Psi} D\!\!\!/ \, {}^2 \Psi$$

Motivation
The Matrix Model
Scalars

# Gravity

Unimodular Met
Seeley-de Witt

EH Actio

UV/IR mixin

Conclusions

# **Emergent Gravity**

### Unimodular metric

$$\widetilde{G}_{ab}:=e^{\sigma}G_{ab} \qquad e^{\sigma} = (\det G_{ab})^{-1/4} 
ightarrow \det \widetilde{G}=1$$
  $S_{ ext{square}}=\int d^4y \sqrt{\widetilde{G}}\ ar{\varPsi} \widetilde{D}^2 \Psi$ 

$$\widetilde{\mathcal{D}}^{2} \Psi = -\left(\widetilde{G}^{ab} \partial_{a} \partial_{b} \Psi + e^{-\sigma} \gamma_{a} \gamma_{b} \theta^{ma} \left(\partial_{m} \theta^{db}\right) \partial_{d} \Psi\right) 
:= -\left(\widetilde{G}^{ab} \partial_{a} \partial_{b} \Psi + \widetilde{a}^{d} \partial_{d} \Psi\right)$$

Motivation
The Matrix Model
Scalars

Emergent Gravity

Upon Quantizatio
Unimodular Metri
Seeley-de Witt
coefficients

EH Actio

UV/IR mixin

Conclusions

# Seeley-de Witt coefficients for Fermions

$$\frac{1}{2}\mathrm{Tr}\Big(\log\widetilde{\mathcal{D}}^2 - \log\widetilde{\mathcal{D}}_0^2\Big) = -\frac{1}{2}\mathrm{Tr}\int_0^\infty \frac{d\alpha}{\alpha}\Big(e^{-\alpha\,\widetilde{\mathcal{D}}^2} - e^{-\alpha\,\widetilde{\mathcal{D}}_0^2}\Big)e^{-\frac{1}{2\alpha\,\widetilde{\lambda}^2}}$$

Heat kernel expansion

$$\operatorname{Tr} e^{-\alpha \widetilde{\mathcal{D}}^2} = \sum_{n \geq 0} \alpha^{\frac{n-4}{2}} \int_{\mathcal{M}} d^4 y \, a_n \left( y, \widetilde{\mathcal{D}}^2 \right)$$

Seeley-de Witt coefficients

$$a_0(y) = rac{1}{16\pi^2} \operatorname{tr} \mathbb{I}$$
  $a_2(y) = rac{1}{16\pi^2} \operatorname{tr} \left( rac{R[\widetilde{G}]}{6} \mathbb{I} + \mathcal{E} 
ight)$ 

Motivation
The Matrix Model
Scalars

### Emergent Gravity

Upon Quantizatio
Unimodular Metr
Seeley-de Witt

EH Actio

UV/IR mixin

Conclusions

# Effective Action for Fermions

$$\Gamma_{\Psi} = \frac{1}{16 \pi^2} \int d^4y \left( 2 \operatorname{tr}(\mathbb{I}) \widetilde{A}^4 + \operatorname{tr} \left( \frac{R[\widetilde{G}]}{6} \mathbb{I} + \mathcal{E} \right) \widetilde{A}^2 + O(\log \widetilde{A}) \right)$$

Commutative case:

$$tr\mathcal{E}_{comm} = -R$$

This model: no spin connection  $\rightarrow \mathcal{E}$  will be modified Still a reasonable Einstein-Hilbert action?

yes

Motivation
The Matrix Model
Scalars

Emergent Gravity

Upon Quantizatio
Unimodular Metr
Seeley-de Witt

EH Action

UV/IR mixing

# Induced EH-Action?

$$\mathcal{E} = -\widetilde{G}^{mn} \left( \partial_m \Omega_n + \Omega_m \Omega_n - \widetilde{\Gamma}_{mn}^k \Omega_k \right)$$

$$\Omega_m = \frac{1}{2} \, \widetilde{G}_{mn} \left( \widetilde{a}^n + \widetilde{\Gamma}^n \right) \qquad \widetilde{a}^n = e^{-\sigma} \gamma_a \gamma_b \theta^{ma} \left( \partial_m \theta^{nb} \right)$$

$$\int d^4 y \operatorname{tr} \mathcal{E} = - \int d^4 y \left( 2R[\widetilde{G}] - G^{mn} \left( \partial_m \sigma \right) \left( \partial_n \sigma \right) \right)$$

for on-shell geometries ↔ fulfill eom

Obtain EH action with an unsual numerical constant

+ dilaton-like term

### EH Action

UV/IR mixing

Conclusions

# $R \& tr \mathcal{E}$

# Use Jacobi identity

$$\partial_a \theta_{bc}^{-1} + \partial_c \theta_{ab}^{-1} + \partial_b \theta_{ca}^{-1} = 0$$

$$\begin{split} R[\widetilde{G}] &= e^{-\sigma} \left[ -\frac{1}{2} \theta^{mn} G^{pq} \partial_p \partial_q \theta_{mn}^{-1} - \frac{1}{2} G^{pq} \left( \partial_p \theta^{mn} \right) \left( \partial_q \theta_{mn}^{-1} \right) \right. \\ & \left. - \left( \partial_m \theta^{ma} \right) G^{nk} \left( \partial_k \theta_{na}^{-1} \right) + \frac{1}{2} G^{mn} \left( \partial_m \sigma \right) \left( \partial_n \sigma \right) \right. \\ & \left. + \frac{1}{2} \left( \partial^m \theta_{na}^{-1} \right) \left( \partial^n \theta_{mb}^{-1} \right) g^{ab} - \frac{1}{2} G^{mn} \left( \partial^q \theta_{ma}^{-1} \right) \left( \partial_q \theta_{nb}^{-1} \right) g^{ab} \right. \\ & \left. - \frac{1}{2} \left( \partial_m \theta^{na} \right) \left( \partial_n \theta^{mb} \right) g_{ab} \right] \\ \mathrm{tr} \mathcal{E} &= e^{-\sigma} \left[ G^{mn} \left( \partial^l \theta_{ma}^{-1} \right) \left( \partial_l \theta_{nb}^{-1} \right) g^{ab} - \left( \partial^m \theta_{na}^{-1} \right) \left( \partial^n \theta_{mb}^{-1} \right) g^{ab} \right] \end{split}$$

# Gravity Upon Quantization

Upon Quantizatio
Unimodular Meta
Seeley-de Witt

### EH Action

UV/IR mixing

Conclusions

# Partial integration reduces Ricci scalar to

$$\int d^4y R[\widetilde{G}]\widetilde{A}^2 = e^{-\sigma} \times$$

$$\left[ \frac{1}{2} \left( \partial^m \theta_{na}^{-1} \right) \left( \partial^n \theta_{mb}^{-1} \right) g^{ab} - \frac{1}{2} G^{mn} \left( \partial^p \theta_{ma}^{-1} \right) \left( \partial_p \theta_{nb}^{-1} \right) g^{ab} \right.$$

$$\left. - \frac{1}{2} \left( \partial_p \theta^{pa} \right) G^{qk} \left( \partial_k \theta^{-1} \right) + \frac{1}{2} G^{mn} \left( \partial_m \sigma \right) \left( \partial_n \sigma \right) \right] \widetilde{A}^2$$

EH Action

U V/IK IIIIXII

Conclusion

### Cancellation of bosonic and fermionic contributions

$$\Gamma_{\Psi} + 4\Gamma_{\Phi} = rac{1}{16\pi^2} \int d^4y \, \mathrm{tr} \, \mathcal{E} \, \widetilde{\Lambda}^2 \, .$$

- Induced gravitational action is nontrivial in the case of e.g.
   N = 1 supersymmetry
- UV/IR mixing remains in supersymmetric models
- Full cancellation only for N = 4 supersymmetry

### UV/IR mixing

# Relation with gauge theory ...

$$S = (2\pi)^2 \operatorname{Tr} \bar{\Psi} \gamma_a \left[ Y^a, \Psi \right]$$

regarded as action for fermions on  $\mathbb{R}^{\frac{4}{a}}$  coupled to U(1) gauge field

covariant coordinates

$$Y^a = X^a - \bar{\theta}^{ab} A_b(x)$$

 $A_b(x)$  hermitian matrices, smooth functions on  $\mathbb{R}^4_{\bar{a}}$ 

$$\left[X^a, X^b\right] = i\,\bar{\theta}^{ab}$$

$$\bar{\theta}^{ab}$$
 constant

$$\bar{g}^{ab} = \bar{\theta}^{am}\bar{\theta}^{bn}g_{mn} \qquad \bar{\rho} = |\bar{g}_{ab}|^{1/4}$$

$$\bar{\rho} = |\bar{g}_{ab}|^{1/4}$$

Motivation
The Matrix Model
Scalars

### Emergent Gravity

Upon Quantizatio Unimodular Meta Seeley-de Witt coefficients

EH Actio

### UV/IR mixing

Conclusions

# UV/IR mixing

The "conventional" gauge theory point of view.

Write the action

$$S = (2\pi)^2 \operatorname{Tr} \bar{\Psi} \gamma_a \left[ Y^a, \Psi \right]$$

as NC  $\overline{U(1)}$  gauge theory...

$$S = \int d^4x \, \bar{\Psi} i \gamma^a \, \left( \bar{\partial}_a \Psi + ig \left[ A_a, \Psi 
ight] 
ight)$$

Consider again the one-loop effective action

→ Use fermionic Feynman rules

Motivation
The Matrix Model
Scalars

### Emergent Gravity

Upon Quantizatio
Unimodular Metri
Seeley-de Witt
coefficients

EH Action

### UV/IR mixing

Conclusion

# UV/IR mixing

$$\varGamma_{\varPsi} = -\frac{1}{2} \mathrm{Tr} \log \Delta_0 - \frac{g^2}{2} \langle \int d^4 x \, \bar{\rho} \bar{\varPsi} \, \gamma^a \, [A_a, \varPsi] \int d^4 y \, \bar{\rho} \bar{\varPsi} \, \gamma^b \, [A_b, \varPsi] \rangle$$

... corresponds to the Feynman diagram



$$\Gamma_{\Psi} = -4\Gamma_{\Phi} - \int d^4x \, \bar{\rho} \, \bar{g}^{ac} \bar{g}^{bd} \, \bar{F}_{ab} \bar{\partial}^2 \bar{F}_{cd} \frac{\Lambda^2}{2}$$

Motivation
The Matrix Model
Scalars

### Emergent Gravity

Upon Quantizatio
Unimodular Metr
Seeley-de Witt

EH Actio

UV/IR mixing

Conclusions

# Rewrite geometric $\Gamma_{\Psi}$

We had

$$egin{aligned} arGamma_{\Psi} &= rac{1}{16\,\pi^2} \int d^4y \left( 2\, ext{tr}(\mathbb{I})\,\widetilde{A}^4 + ext{tr}\left(rac{R[\widetilde{G}]}{6}\,\mathbb{I} + \mathcal{E}
ight)\widetilde{A}^2 + O(\log\widetilde{A}) 
ight) \ &= -4\,arGamma_{\Phi} + rac{1}{16\pi^2} \int d^4y\, ext{tr}\mathcal{E}\widetilde{A}^2 \end{aligned}$$

compare this to the result from gauge theory point of view

rewrite  $\Gamma_{\Psi}$  in x-coordinates of Moyal-Weyl plane

Motivation
The Matrix Model
Scalars

### Emergent Gravity

Upon Quantizati Unimodular Met Seeley-de Witt coefficients

EH Actio

UV/IR mixing

Conclusions

# UV/IR mixing ... ...an effect of gravity

Relation between y and x coordinates

$$y^a = x^a - \bar{\theta}^{ab}\bar{A}_b + O(\theta^2)$$

Relation between

$$i \theta^{ab}(y) = \left[ Y^a, Y^b \right] = i \bar{\theta}^{ab} - i \bar{\theta}^{ac} \bar{\theta}^{bd} \bar{F}_{cd}$$
  
 $\theta_{ab}^{-1}(y) = \bar{\theta}_{ab}^{-1} - \bar{F}_{ab}$ 

Obtain effective action in x-coordinates

$$\Gamma_{\Psi} = -4\Gamma_{\Phi} - \int d^4x \,\bar{\rho} \,\bar{g}^{ac} \bar{g}^{bd} \,\bar{F}_{ab} \bar{\partial}^2 \bar{F}_{cd} \frac{\Lambda^2}{2}$$

UV/IR mixing is understood as an effect of gravity.

Motivation
The Matrix Model
Scalars
Fermions

### Emergent Gravity

Upon Quantization
Unimodular Meta
Seeley-de Witt

EH Action

JV/IR mixin

Conclusions

# Conclusions

Framework of NC emergent gravity extends to fermions

Fermions couple to background geometry

Spin connection appears to be missing

Fermions will follow standard trajectories with different spin rotation

Einstein-Hilbert action is induced

UV/IR mixing is an effect of gravity

Bosonic & fermionic contributions do not cancel

... only for N = 4 SUSY