

# Field theory on curved NC spacetimes

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Bayrischzell Workshop 2010

*Noncommutativity and Physics:  
Spacetime Quantum Geometry*

Bayrischzell, Germany, May 14-17, 2010



# Motivation

- ▶ QFT on curved spacetimes is important for physics

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  - ▶  $\exists$  NC gravity solutions [[Schupp, Solodukhin; TO, AS; Aschieri, Castellani](#)]
  - test their physical implications by using QFTCS

# Scalar field theory on a class of curved NC spacetimes

- ▶ *Simple example of a twist:* [Moyal product/twist]

★-product 
$$h \star k = h e^{\frac{i\lambda}{2} \overleftarrow{\partial}_\mu \Theta^{\mu\nu} \overrightarrow{\partial}_\nu} k$$



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- ▶ *Our class of twists:*  $\mathcal{F}^{-1} = \tilde{f}^\alpha \otimes_{\mathbb{C}} \tilde{f}_\alpha \in \text{UVec}[[\lambda]] \otimes_{\mathbb{C}} \text{UVec}[[\lambda]]$

- ▶ normalization:  $(\epsilon \otimes_{\mathbb{C}} \text{id})\mathcal{F} = (\text{id} \otimes_{\mathbb{C}} \epsilon)\mathcal{F} = 1$
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- ▶ algebra of functions  $(C^\infty(\mathcal{M})[[\lambda]], \star)$ , where  $\mathfrak{h} \star \mathfrak{k} := \bar{f}^\alpha(\mathfrak{h}) \cdot \bar{f}_\alpha(\mathfrak{k})$

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- ▶ pairing  $\langle \nu, \omega \rangle_\star := \langle \bar{f}^\alpha(\nu), \bar{f}_\alpha(\omega) \rangle$  among vflds and 1 forms

- ▶ Action (basis independent):

$$S_\Phi = \int L_\Phi = -\frac{1}{2} \int (\langle \langle d\Phi, g_\star^{-1\star} \rangle_\star, d\Phi \rangle_\star + M^2 \Phi \star \Phi) \star \text{vol}_\star$$

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$$P_{\star}[\Phi] \star \text{vol}_{\star} :=$$

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**NB:**  $P_\star$  is formally self adjoint w.r.t.  $SP (\varphi, \psi)_\star = \int \varphi^\star \star \psi \star \text{vol}_\star$ , i.e.

$$(\varphi, P_\star[\psi])_\star = (P_\star[\varphi], \psi)_\star$$

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very rough estimate:  $\lambda H \approx t_{\text{pl}} H_{\text{today}} \approx 10^{-60}$

$$2.) \mathcal{F}^{-1} = \exp\left(\frac{i\lambda}{2}(x^i \partial_i \otimes_{\mathbb{C}} \partial_\varphi - \partial_\varphi \otimes_{\mathbb{C}} x^i \partial_i)\right) \Rightarrow e^{i\varphi} \star r = e^\lambda r \star e^{i\varphi}$$

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- ▶ slice of de Sitter space:  $ds^2 = -dt^2 + e^{2Ht} (dx^2 + dy^2 + dz^2)$
- ▶ the following NC spacetimes solve NC Einstein equations [TO, AS]

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- ▶ spherical wave  $\Phi = \phi(t) j_l(kr) Y_{lm}(\theta, \varphi)$  and for simplicity  $a(t) \equiv 1$

$$\partial_t^2 \phi(t) + \left(M^2 + k^2 \frac{e^{\lambda m} + e^{-4\lambda m}}{1 + e^{-3\lambda m}}\right) \phi(t) = 0, \quad E(k, m)^2 \geq M^2 + k^2$$

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- ▶ UV improvement of  $\Phi^4$ -theory in  $d = 4$ :

$$\text{loop} \sim \sqrt{\Lambda_{\text{UV}}} + \text{finite}, \quad \text{cross} = \text{finite}$$

Let's go back to curved NC spacetimes and talk about the quantization of scalar fields.

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- ▶ based on strong results for the commutative case we find:

there exist unique Green's operators  $\Delta_{\star\pm} := \sum \lambda^n \Delta_{(n)\pm}$  satisfying

- (i)  $P_\star \circ \Delta_{\star\pm} = \text{id}_{C_0^\infty(\mathcal{M})[[\lambda]]}$ ,
- (ii)  $\Delta_{\star\pm} \circ P_\star|_{C_0^\infty(\mathcal{M})[[\lambda]]} = \text{id}_{C_0^\infty(\mathcal{M})[[\lambda]]}$ ,
- (iii)  $\text{supp}(\Delta_{(n)\pm}[\varphi]) \subseteq J_\pm(\text{supp}(\varphi))$ , for all  $n \in \mathbb{N}^0$  and  $\varphi \in C_0^\infty(\mathcal{M})$ ,

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**NB:** also possible for deformed normally hyperbolic operators  $P_\star$

**Explicit formula** for  $\Delta_{*\pm}$  in terms of  $\Delta_{\pm} := \Delta_{(0)\pm}$ :

$$\begin{aligned} \Delta_{*\pm} &= \Delta_{\pm} \\ &\quad - \lambda \Delta_{\pm} \circ P_{(1)} \circ \Delta_{\pm} \\ &\quad - \lambda^2 (\Delta_{\pm} \circ P_{(2)} \circ \Delta_{\pm} - \Delta_{\pm} \circ P_{(1)} \circ \Delta_{\pm} \circ P_{(1)} \circ \Delta_{\pm}) \\ &\quad + \mathcal{O}(\lambda^3) \quad [\text{higher orders follow the same structure}] \end{aligned}$$

Graphically:

$$\begin{aligned} \equiv &= \text{---} - \lambda \text{---} \textcircled{1} \text{---} - \lambda^2 \left( \text{---} \textcircled{2} \text{---} - \text{---} \textcircled{1} \textcircled{1} \text{---} \right) \\ &- \lambda^3 \left( \text{---} \textcircled{3} \text{---} - \text{---} \textcircled{1} \textcircled{2} \text{---} - \text{---} \textcircled{2} \textcircled{1} \text{---} + \text{---} \textcircled{1} \textcircled{1} \textcircled{1} \text{---} \right) + \mathcal{O}(\lambda^4) \end{aligned}$$

→ perturbative approach to deformed Green's operators

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### Proposition (TO, AS)

$(V_\star, \omega_\star)$  with  $V_\star := H/\text{Ker}(\Delta_\star)$  and

$$\omega_\star([\varphi], [\psi]) := (\varphi, \Delta_\star[\psi])_\star = \int \varphi^* \star \Delta_\star[\psi] \star \text{vol}_\star$$

is a symplectic vector space.

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**NB:** twisting requires CCR-compatible  $\mathcal{F}$  [[Poisson geometry: Aschieri, Lizzi, Vitale](#)]

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- ▶ **Outlook** and future work:

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  - ▶ existence, uniqueness and construction of the deformed Green's operators
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