The covariant derivative

paul.romatschke@colorado.edu

Spring 2021

• In lecture 10, we derived the geodesic equation

$$\frac{d^2 x^{\mu}}{d\tau^2} + \Gamma^{\mu}_{\alpha\beta} \frac{dx^{\alpha}}{d\tau} \frac{dx^{\beta}}{d\tau} = 0.$$
 (16.1)

- We discussed that $\frac{d^2 x^{\mu}}{d \tau^2}$ is not a vector under coordinate transformations
- In this lecture, we will define a new operation that *does* transform properly under coordinate transformations

Vectors and Non-Vectors

• Recall that we are interested in non-linear coordinate transformations

$$dx^{\mu} \to dx^{\prime \mu} = R^{\mu}_{\ \nu} dx^{\nu} \,, \tag{16.2}$$

where $R^{\mu}_{\ \nu}$ does depend on the coordinates

- In the geodesic equation, we have the Christoffel symbols $\Gamma^{\mu}_{\alpha\beta}$
- We may ask ourselves how these transform under (16.2)
- A somewhat tedious exercise shows that

$$\Gamma^{\prime\mu}_{\alpha\beta} = R^{\mu}_{\ \mu'}(R^{-1})^{\ \alpha'}_{\alpha}(R^{-1})^{\ \beta'}_{\beta}\Gamma^{\mu'}_{\alpha'\beta'} - \frac{\partial x^{\prime\mu}}{\partial x^{\alpha'}\partial x^{\beta'}}\frac{\partial x^{\alpha'}}{\partial x^{\alpha}}\frac{\partial x^{\beta'}}{\partial x^{\beta}}.$$
 (16.3)

• Clearly,
$$\Gamma^{\mu}_{\alpha\beta}$$
 is **not** a tensor under (16.2)

Vectors under General Coordinate Transformations

- If A^μ is a vector under (16.2), ∂_νA^μ will in general not be a vector because of coordinate dependence of R^μ_ν
- In more detail, using $A^{\mu} \rightarrow A'^{\mu} = R^{\mu}_{\ \alpha} A^{\alpha}$ we have

$$\partial_{\nu}A^{\mu} \to \partial_{\nu}'A'^{\mu} = \partial_{\nu}'\left(R^{\mu}_{\ \alpha}A^{\alpha}\right) = R^{\mu}_{\ \alpha}\partial_{\nu}'A^{\alpha} + A^{\alpha}\partial_{\nu}'R^{\mu}_{\ \alpha}.$$
(16.4)

• The second term spoils general covariance of $\partial_{\nu}A^{\mu}$

Repairing Covariance

• Very explicitly, using $R^{\mu}_{\ lpha}=rac{\partial x'^{\mu}}{\partial x^{lpha}}$ we have

$$\partial_{\nu}^{\prime} A^{\prime \mu} = R^{\mu}_{\ \alpha} \partial_{\nu}^{\prime} A^{\alpha} + A^{\alpha} \frac{\partial x^{\prime \mu}}{\partial x^{\alpha} \partial x^{\beta}} \frac{\partial x^{\beta}}{\partial x^{\prime \nu}}$$
(16.5)

- The offending second piece looks a lot like the transformation of the Christoffels, cf. (16.3)
- Let's consider the combination

$$\partial_{\nu}A^{\mu} + \Gamma^{\mu}_{\nu\rho}A^{\rho} \equiv \nabla_{\nu}A^{\mu} \,. \tag{16.6}$$

• Using the transformation properties of both $\Gamma^{\mu}_{\nu\rho}$ and A^{μ} , it is straightforward to show that

$$\nabla_{\nu} A^{\mu} \to \nabla_{\nu}^{\prime} A^{\prime \mu} = R^{\mu}_{\alpha} (R^{-1})^{\beta}_{\nu} \nabla_{\beta} A^{\alpha} .$$
(16.7)

The Covariant Derivative

- One can generalize this idea to arbitrary rank tensors
- Defining

$$\nabla_{\nu} X^{\mu_{1}\mu_{2}...\mu_{n}}_{\ \rho_{1}\rho_{2}...\rho_{m}} = \partial_{\nu} X^{\mu_{1}\mu_{2}...\mu_{n}}_{\ \rho_{1}\rho_{2}...\rho_{m}} + \Gamma^{\mu_{1}}_{\nu\alpha} X^{\alpha\mu_{2}...\mu_{n}}_{\ \rho_{1}\rho_{2}...\rho_{m}} + \Gamma^{\mu_{n}}_{\nu\alpha} X^{\mu_{1}\mu_{2}...\mu_{n}}_{\ \rho_{1}\rho_{2}...\rho_{m}} - \Gamma^{\beta}_{\nu\rho_{1}} X^{\mu_{1}\mu_{2}...\mu_{n}}_{\ \beta\rho_{2}...\rho_{m}} - \dots - \Gamma^{\beta}_{\nu\rho_{m}} X^{\mu_{1}\mu_{2}...\mu_{n}}_{\ \rho_{1}\rho_{2}...\beta}$$

these combinations transform as tensors under (16.2)

 We call ∇_µ the (geometric-) covariant derivative and will employ this symbol in order to distinguish it from the usual derivative (∂_µ)

Examples of The Covariant Derivative

Examples for the covariant derivative include

Scalars:

$$\nabla_{\nu} X = \partial_{\nu} X \,. \tag{16.8}$$

Vectors:

$$\nabla_{\nu} X^{\mu} = \partial_{\nu} X^{\mu} + \Gamma^{\mu}_{\nu\alpha} X^{\alpha} \,. \tag{16.9}$$

• Rank 2 tensors:

$$\nabla_{\nu} X^{\alpha\beta} = \partial_{\nu} X^{\alpha\beta} + \Gamma^{\alpha}_{\nu\rho} X^{\rho\beta} + \Gamma^{\beta}_{\nu\rho} X^{\alpha\rho}$$
(16.10)

Importance of Covariant Derivative

- Covariant derivative generalizes "flat-space" identities
- For instance, $\partial_{\mu}j^{\mu} = 0$ is **not** covariant under (16.2)
- Instead of $\partial_{\mu}j^{\mu} = 0$, the correct current conservation law is

$$abla_{\mu} j^{\mu} = 0.$$
 (16.11)

This has important consequences. For instance, on a curved manifold

$$\partial_{\mu}j^{\mu} = -\Gamma^{\mu}_{\mu\alpha}j^{\alpha} \neq 0$$
 (16.12)

(sort of like Coriolis force: not "real force", but you'll feel it anyway!)



Importance of Covariant Derivative

• The same holds true for energy-momentum conservation. We have

$$\nabla_{\mu}T^{\mu\nu} = 0 = \partial_{\mu}T^{\mu\nu} + \Gamma^{\mu}_{\mu\alpha}T^{\alpha\nu} + \Gamma^{\nu}_{\mu\alpha}T^{\mu\alpha}, \qquad (16.13)$$

which in general implies $\partial_{\mu}T^{\mu\nu} \neq 0$

Specifically

$$\partial_{\mu}T^{\mu\nu} = -\Gamma^{\mu}_{\mu\alpha}T^{\alpha\nu} - \Gamma^{\nu}_{\mu\alpha}T^{\mu\alpha}, \qquad (16.14)$$

so our notion of energy-conservation as in Minkowski space-time no longer holds in GR

 Defining "energy" can become tricky in dynamical situations involving gravity!