

# Our Universe

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## Review

- In lecture 37, we derived a balance equation for the geometry of the universe
- Using the notation  $\Omega \equiv \frac{8\pi G\epsilon}{3H^2}$  we found

$$\frac{\kappa}{a^2 H^2} = \Omega - 1. \quad (38.1)$$

- For an equation of state  $p = w\epsilon$ , we found for the energy density

$$\epsilon(t) = \epsilon_0 \left( \frac{a_0}{a} \right)^{3(1+w)} \quad (38.2)$$

- In lecture 37, we found a solution for  $a(t)$  assuming  $\kappa = 0$  (flat universe)
- In this lecture, we will explore what happens for  $\kappa \neq 0$

## Review – Geometry

- Recall that the value of  $\kappa$  corresponds to the overall geometry of space
- For  $\kappa = 0$ , we have flat space,  $R = 0$  and  $\Lambda = 0$
- For  $\kappa = 1$ , we have elliptic space,  $R > 0$  and  $\Lambda > 0$
- For  $\kappa = -1$ , we have hyperbolic space,  $R < 0$  and  $\Lambda < 0$

# Geometry of the Universe

- We considered the case of  $\kappa = 0$  (flat universe) in lecture 37
- From (38.1) we have  $\kappa = 0 \leftrightarrow \Omega = 1$
- Now for  $\kappa < 0$ , (38.1) implies  $\Omega < 1$ ; we have an **under-dense** universe
- By contrast for  $\kappa > 0$ , (38.1) implies  $\Omega > 1$ : we have an **over-dense** universe

## Energy Density and Geometry

- Depending on value of  $\Omega$ , we have a different geometry of the universe
- Knowing  $\Omega$  is important!
- The value of  $\Omega$  is given by the sum over all matter contributions:

$$\Omega = \Omega_M + \Omega_R + \Omega_\Lambda \quad (38.3)$$

- Each of these contributions comes with a different EoS parameter:

$$\text{Matter : } w = 0, \quad \text{Radiation : } w = \frac{1}{3}, \quad \text{CC : } w = -1. \quad (38.4)$$

- As a consequence, the energy-densities (38.2) scale differently:

$$\epsilon_M \propto a^{-3}, \quad \epsilon_R \propto a^{-4}, \quad \epsilon_\Lambda \propto a^0. \quad (38.5)$$

- In addition, lhs of (38.1) scales as  $\epsilon_\kappa \propto a^{-2}$ ; **Complicated scale factor evolution!**

# Our Universe

We have observational data for our universe:

- $\Omega \simeq 1 = \Omega_M + \Omega_R + \Omega_\Lambda$ ; **our universe is almost flat!**
- $\Omega_{M,0} \simeq 0.3 \pm 0.1$ ; **about 1/3 is matter**
- $\Omega_{R,0} \simeq 0$ ; **almost no radiation**
- As a consequence, we must have  $\Omega_{\Lambda,0} \simeq 0.7$ ; **the lion's share of  $\Omega$  is vacuum energy!**

## More Details: Radiation

- Today, most radiation in the universe is from the Cosmic Microwave Background (CMB)
- This radiation permeates the universe, and it's almost black-body with  $T \simeq 2.7$  K
- The energy density for radiation is

$$\epsilon_R = \frac{\pi^2 T^4}{90} \times 2 \quad (38.6)$$

- The critical energy density today is

$$\epsilon_c = \frac{3H_0^2}{8\pi G} \simeq 5 \text{ GeV} m^{-3}. \quad (38.7)$$

- Using unit conversion, this gives

$$\Omega_{R0} = \frac{\epsilon_R}{\epsilon_c} \simeq 2 \times 10^{-5} \quad (38.8)$$

## More Details: Matter

- Observational data suggests that

$$\Omega_{M0} \simeq 0.3 \quad (38.9)$$

- **Normal** matter is mostly in stars; for stars, we can calculate

$$\Omega_{\text{stars}} \simeq 0.046 . \quad (38.10)$$

- This leaves a gap of

$$\Omega_{M0} - \Omega_{\text{stars}} \simeq 0.25 . \quad (38.11)$$

- **We have no idea what this matter is!**
- We call it “Dark Matter”



## More Details: Dark Energy

- From the energy budget (38.3), we found

$$\Omega_{\Lambda 0} \simeq 0.7 \quad (38.12)$$

- **We cannot calculate  $\Omega_{\Lambda}$ !**
- A naive estimate gives  $\Omega_{\Lambda 0} \simeq 10^{120}$ , which is quite a bit off
- We call  $\Omega_{\Lambda}$  **dark energy** (mostly because it sounds mysterious)
- We **could** calculate  $\Omega_{\Lambda}$  if we understood how to consistently formulate a **quantum theory of gravity**

# Coincidence problem

- Observational data suggests

$$\Omega_{R0} \propto 10^{-5}, \quad \Omega_{M0} \propto 0.3, \quad \Omega_{\Lambda 0} \propto 0.7. \quad (38.13)$$

- We know that

$$\Omega_R \propto a^{-4}, \quad \Omega_M \propto a^{-3}, \quad \Omega_{\Lambda} \propto a^0. \quad (38.14)$$

- Over the history of the universe,  $a(t)$  changes dramatically
- Why is  $\Omega_M \propto \Omega_{\Lambda}$  **today**?
- This is known as the **coincidence problem**