## Regularizing QFT Divergencies

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#### References

• Laine and Vuorinen, "Basics of Thermal Field Theory", Chapter 2:

► https://arxiv.org/pdf/1701.01554.pdf

## A Divergence in the free QFT partition function

Recall from Lecture 9 that the free QFT partition function is given by

$$\ln Z_{\rm free} = -\frac{V}{T} \int \frac{d^D k}{(2\pi)^D} \left( \frac{E_k}{2} + T \ln \left( 1 - e^{-\beta E_k} \right) \right) , \qquad (10.1)$$

where D denotes the number of space dimensions and  $E_k = \sqrt{\vec{k}^2 + m^2}$ .

- Instead of  $Z_{\text{free}}$ , let us consider a physically intuitive observable, the pressure p(T)
- ullet Given any partition function Z, basic thermodynamic relations define the pressure as

$$p(T) = \frac{T}{V} \ln Z. \tag{10.2}$$

## The pressure of free scalar quantum field theory

• From the partition function (10.1), the pressure of a single free scalar field is

$$p(T) = -\int \frac{d^D k}{(2\pi)^D} \left( \frac{E_k}{2} + T \ln \left( 1 - e^{-\beta E_k} \right) \right). \tag{10.3}$$

- This expression for the pressure is divergent for all temperatures T
- In particular, the divergence is there also for zero temperature

$$p(0) = -\frac{1}{2} \int \frac{d^D k}{(2\pi)^D} E_k = -\frac{1}{2} \int \frac{d^D k}{(2\pi)^D} \sqrt{\vec{k}^2 + m^2}$$
 (10.4)

#### Regulating a divergent function

• For further convenience, let us study the slightly more general function

$$\Phi(m, D, A) = \int \frac{d^D k}{(2\pi)^D} \left(\vec{k}^2 + m^2\right)^{-A} . \tag{10.5}$$

• In terms of this function, the zero-temperature free scalar field theory pressure in 3 space dimensions is given by

$$\rho(0) = -\frac{1}{2}\Phi\left(m, 3, -\frac{1}{2}\right). \tag{10.6}$$

## Regulating a divergent function

• Looking at the integrand of (10.5), we see that it only depends on  $|\vec{k}| \equiv k$ , so we can do the angular integral

$$\Phi(m, D, A) = \frac{\Omega_D}{(2\pi)^D} \int_0^\infty dk k^{D-1} \left(k^2 + m^2\right)^{-A}, \qquad (10.7)$$

where  $\Omega_D = \frac{2\pi^{\frac{D}{2}}}{\Gamma(\frac{D}{2})}$  is the "solid angle" in D spatial dimensions, e.g.  $\Omega_3 = 4\pi$ ,  $\Omega_2 = 2\pi$ , etc.

- The integrand of (10.7) is well-behaved for any  $k \in [0, \infty)$  as long as  $m \neq 0$
- Assuming  $m \neq 0$  in the following, the only divergence for  $\Phi$  arises for high wave-number  $k \to \infty$
- Borrowing from the nomenclature of light at high wavenumber, this is called an "ultraviolet" (UV) divergence
- Later on we will also encounter divergencies for  $k \to 0$ , which are called "infrared" (IR) divergencies

## Regulating UV divergencies

- ullet Let us first study the degree of the UV divergence of  $\Phi$
- Since the divergence is in the UV, we can concentrate on the high momentum modes k, and neglect the mass scale in the integrand
- $\bullet$  Then we cut the momentum integral off at the scale  $\Lambda\gg 1$  and find the degree of divergence as

$$\int_0^{\Lambda} dk k^{D-1} \left( k^2 + m^2 \right)^{-A} \propto \int_0^{\Lambda} dk k^{D-1} k^{-2A} \propto \frac{\Lambda^{D-2A}}{D - 2A}, \quad (10.8)$$

- In case of the pressure with D=3,  $A=-\frac{1}{2}$ , this gives  $p(0)\propto \Lambda^4$
- ullet The pressure (10.4) diverges with degree four as  $\Lambda o \infty$
- ullet Actually we can evaluate p(0) exactly using cut-off regularization

$$p(0) = -\frac{m^4}{16\pi^2} \left[ \frac{\Lambda^4}{m^4} + \frac{\Lambda^2}{2m^2} - \frac{1}{2} \ln \left( \frac{2\Lambda}{m} \right) \right]$$
 (10.9)

- Now let us try a different regularization scheme
- Instead of introducing a UV cut-off  $\Lambda$ , we can realize that the integral for  $\Phi$  can be solved exactly:

$$\Phi(m, D, A) = \frac{\Omega_D}{(2\pi)^D} \frac{\Gamma(A - \frac{D}{2})\Gamma(1 + \frac{D}{2})}{D\Gamma(A)} (m^2)^{-A + \frac{D}{2}} (10.10)$$

$$= \frac{1}{(4\pi)^{\frac{D}{2}}} \frac{\Gamma(A - \frac{D}{2})}{\Gamma(A)} (m^2)^{-A + \frac{D}{2}},$$

- The divergence for  $A=-\frac{1}{2}$  and D=3 is apparent from the divergence of the  $\Gamma$  function with negative argument,  $\Gamma(-2)=\infty$
- But what if we work in "almost" three dimensions, e.g.

$$D = 3 - 2\varepsilon, \tag{10.11}$$

with  $\lim \varepsilon \to 0$ ?

• For  $D = 3 - 2\varepsilon$ , we have

$$\Phi(m, 3 - 2\epsilon, -\frac{1}{2}) = \frac{1}{(4\pi)^{\frac{3}{2} - \varepsilon}} \frac{\Gamma(-2 + \varepsilon)}{\Gamma(-\frac{1}{2})} (m^2)^{2 - \varepsilon}, \qquad (10.12)$$

which is finite for any  $\varepsilon \in (0,1)$ .

• We can use the property  $x\Gamma(x)=\Gamma(1+x)$  of the  $\Gamma$ -function to write

$$\Gamma(-2+\varepsilon) = \frac{\Gamma(1+\varepsilon)}{(-2+\varepsilon)(-1+\varepsilon)\varepsilon}$$
 (10.13)

Then, we can expand all components of  $\Phi$  in a power series for small  $\varepsilon$ , e.g.

$$(4\pi)^{-\frac{3}{2}+\varepsilon} = \frac{2\sqrt{\pi}}{(4\pi)^2} \left[ 1 + \varepsilon \ln(4\pi) + \mathcal{O}(\varepsilon^2) \right] ,$$

$$(m^2)^{2-\varepsilon} = m^4 (m^2)^{-\varepsilon} = m^4 \left[ 1 - \varepsilon \ln(m^2) + \mathcal{O}(\varepsilon^2) \right] ,$$

$$\Gamma(-2+\varepsilon) = \frac{\Gamma(1+\varepsilon)}{(-2+\varepsilon)(-1+\varepsilon)\varepsilon} = \frac{1}{2\varepsilon} \left[ 1 + \varepsilon \left( \frac{3}{2} - \gamma_E \right) + \mathcal{O}(\varepsilon^2) \right] ,$$

where  $\gamma_E = 0.577215664901...$  is Euler's constant.

 Putting everything together, we obtain for the zero-temperature pressure

$$p(0) = \frac{m^4}{64\pi^2} \left[ \frac{1}{\varepsilon} - \ln(m^2) + \ln(4\pi) - \gamma_E + \frac{3}{2} + \mathcal{O}(\varepsilon) \right]$$
 (10.15)

- ullet The pressure diverges as arepsilon o 0, as with cut-off regularization
- There's only one weird thing: there is a mass squared under the logarithm; we shouldn't have a dimensionful quantity in the log!
- It's coming from the expansion of  $(m^2)^{2-\varepsilon}$  in (10.14)
- And *this* is coming from the integral (10.5) having non-integer mass dimension for  $D=3-2\varepsilon$

• To avoid a dimensionful quantity under the logarithm, we re-define the function  $\Phi$  from (10.5) in  $D=3-2\varepsilon$  dimensions as

$$\Phi(m, 3-2\varepsilon, A) = \mu^{2\varepsilon} \int \frac{d^{3-2\varepsilon}k}{(2\pi)^{3-2\varepsilon}} \left(\vec{k}^2 + m^2\right)^{-A}. \tag{10.16}$$

- The parameter  $\mu$  is arbitrary, and is taken to have dimensions of energy (mass)
- ullet This ensures that the pressure has mass dimension four even if arepsilon 
  eq 0
- ullet We call  $\mu$  the renormalization scale parameter

• With the redefinition (10.16) for non-integer dimensions, the dimensionally-regulated zero-temperature pressure then reads

$$p(0) = \frac{m^4}{64\pi^2} \left[ \frac{1}{\varepsilon} + \ln\left(\frac{\mu^2}{m^2}\right) + \ln(4\pi) - \gamma_E + \frac{3}{2} + \mathcal{O}(\varepsilon) \right] \quad (10.17)$$

• It is customary to further simplify this expression by introducing instead of  $\mu$  the so-called " $\overline{\rm MS}$  scheme" scale parameter

$$\bar{\mu}^2 = 4\pi \mu^2 e^{-\gamma_E} \tag{10.18}$$

 $\bullet$  In  $\overline{\rm MS}$  then

$$p(0) = \frac{m^4}{64\pi^2} \left[ \frac{1}{\varepsilon} + \ln\left(\frac{\bar{\mu}^2}{m^2}\right) + \frac{3}{2} + \mathcal{O}(\varepsilon) \right]$$
 (10.19)